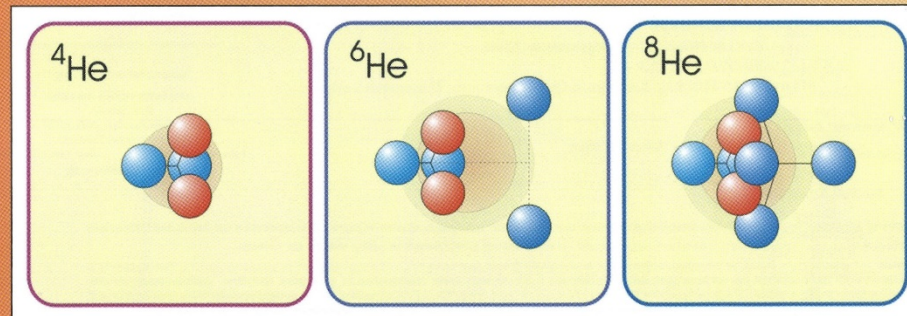


REVIEWS of MODERN PHYSICS™

October–December 2013

Volume 85, Number 4



COLLOQUIUM: LASER PROBING OF NEUTRON-RICH NUCLEI IN LIGHT ATOMS

APS
physics

Published by American Physical Society™



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NATIONAL
LABORATORY

... for a brighter future



U.S. Department
of Energy

UChicago ►
Argonne_{LLC}



A U.S. Department of Energy laboratory
managed by UChicago Argonne, LLC

Simple Atom, Extreme Nucleus: Laser Trapping and Probing of He-8

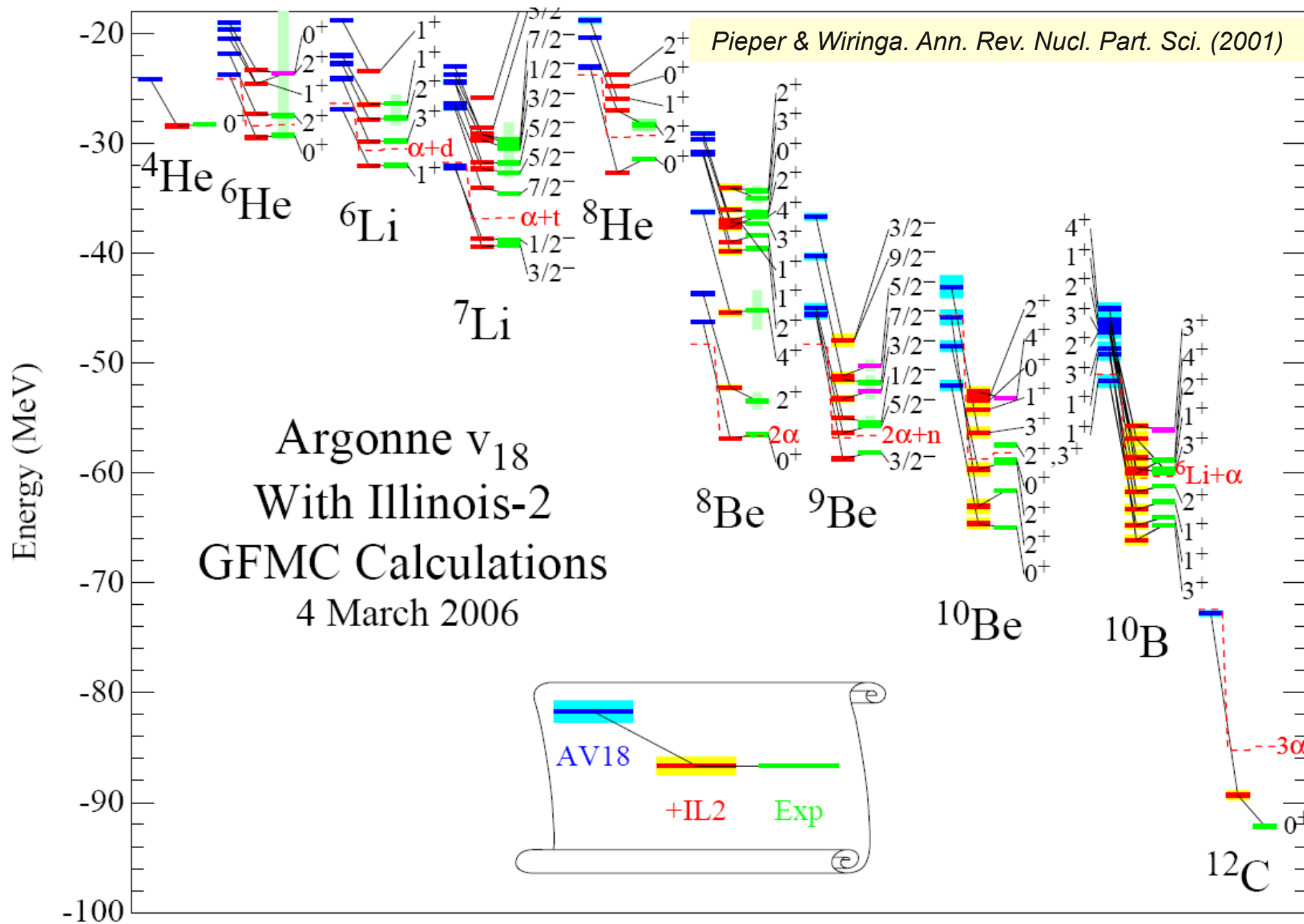
Zheng-Tian Lu

Argonne National Laboratory

University of Chicago

Funding: DOE, Office of Nuclear Physics

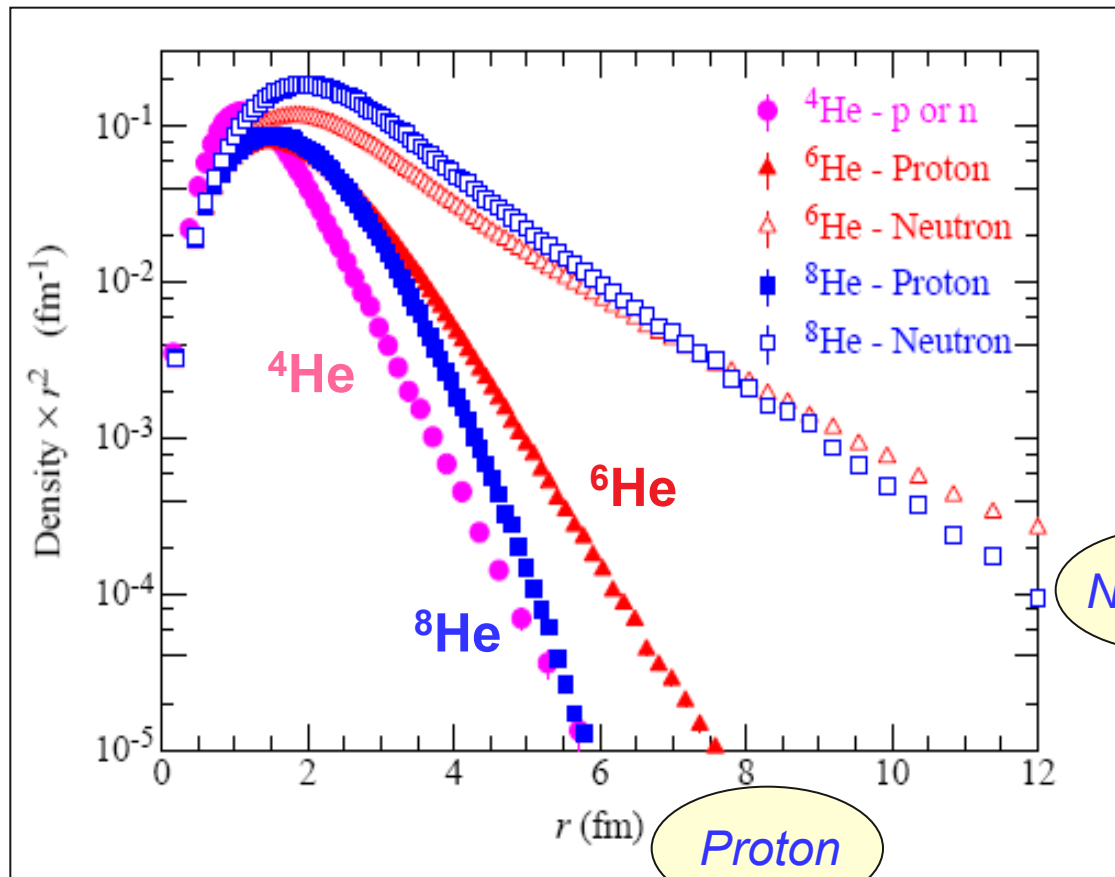
Quantum Monte Carlo Calculations of Light Nuclei



Halo Nuclei ${}^6\text{He}$ and ${}^8\text{He}$

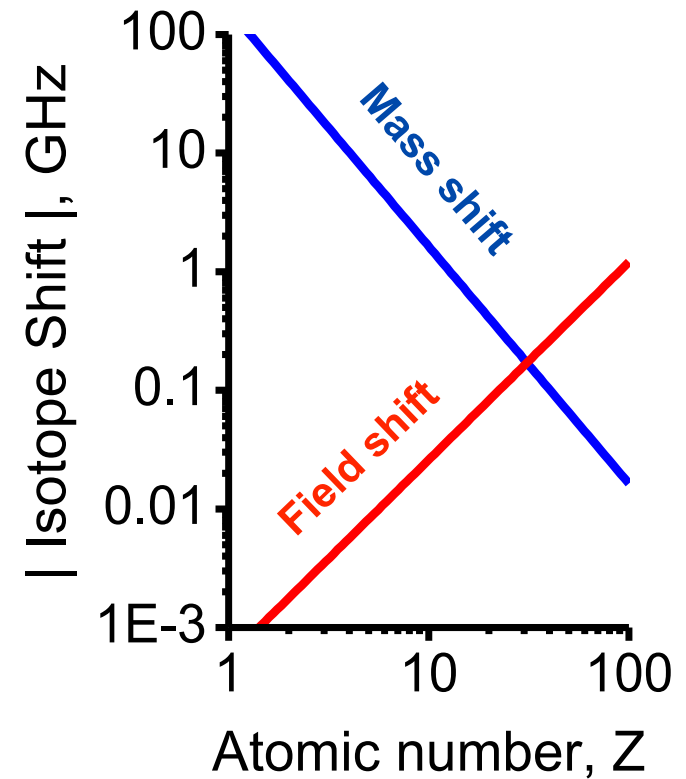
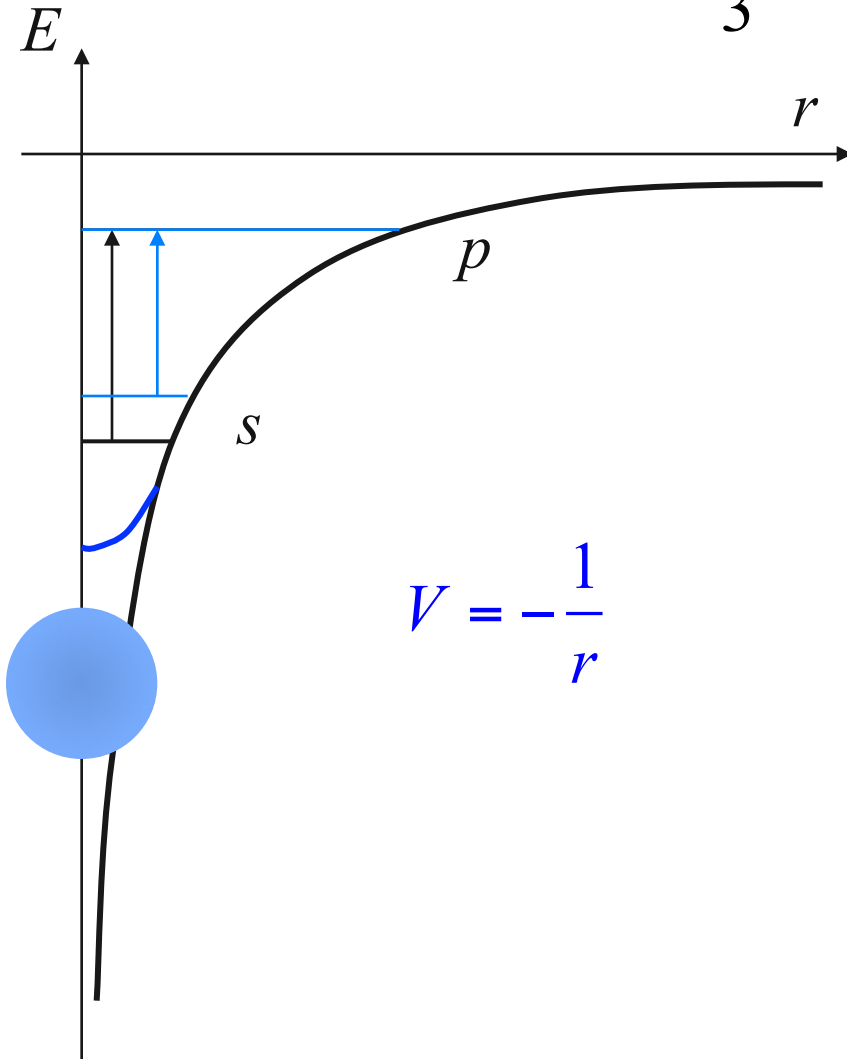
| Isotope | Half-life | Spin | Isospin | Core + Valence |
|---------|-----------|-------|---------|----------------|
| He-6 | 0.8 s | 0^+ | 1 | $\alpha + 2n$ |
| He-8 | 0.1 s | 0^+ | 2 | $\alpha + 4n$ |

Quantum
Monte Carlo
calculation



Nuclear Volume Effect

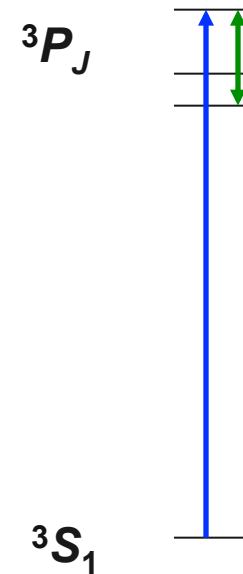
$$\delta v_{FS} = -\frac{2\pi}{3} Ze^2 \cdot \Delta |\Psi(0)|^2 \cdot \delta \langle r^2 \rangle$$



Atomic Theory of Helium

Drake, *Can. J. Phys.* (2006);
Pachucki & Sapirstein, *J. Phys. B* (2002)

- ◆ Perturbation theory with corrections:
 - ◆ relativity
 - ◆ QED
 - ◆ *finite nuclear mass*
 - ◆ *nuclear charge radius*
- ◆ Uncertainty on transition frequency: ~ MHz
- ◆ Uncertainty on isotope shift: (MHz) \times (m_e/M_N) \rightarrow kHz



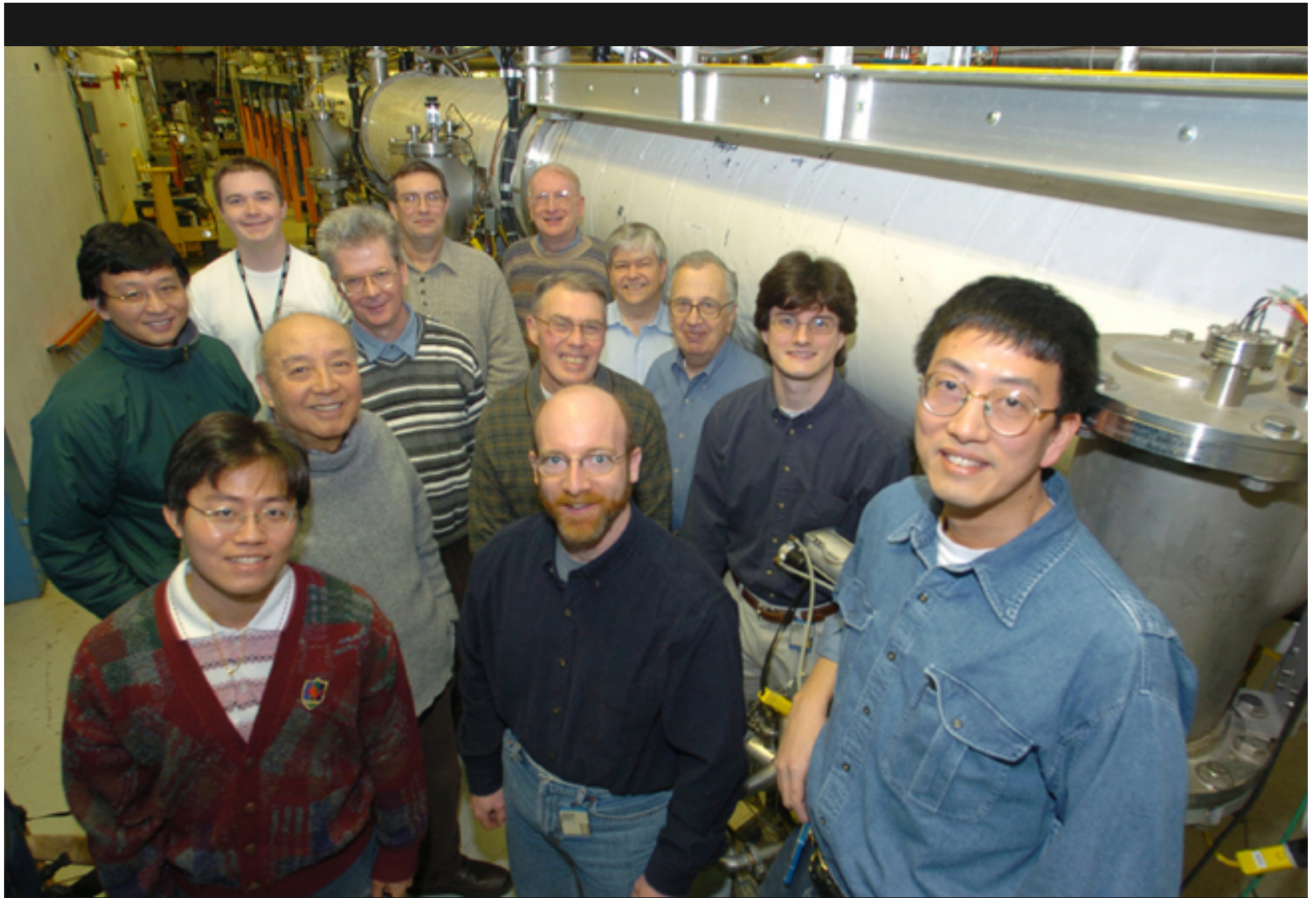
Isotope Shift $dn = dn_{MS} + dn_{FS}$

For $2^3S_1 - 3^3P_2$ transition @ 389 nm:

$${}^6\text{He} - {}^4\text{He} : \delta\nu_{6,4} = 43196.171(2) \text{ MHz} - 1.010 (\langle r^2 \rangle_{c,6} - \langle r^2 \rangle_{c,4}) \text{ MHz/fm}^2$$

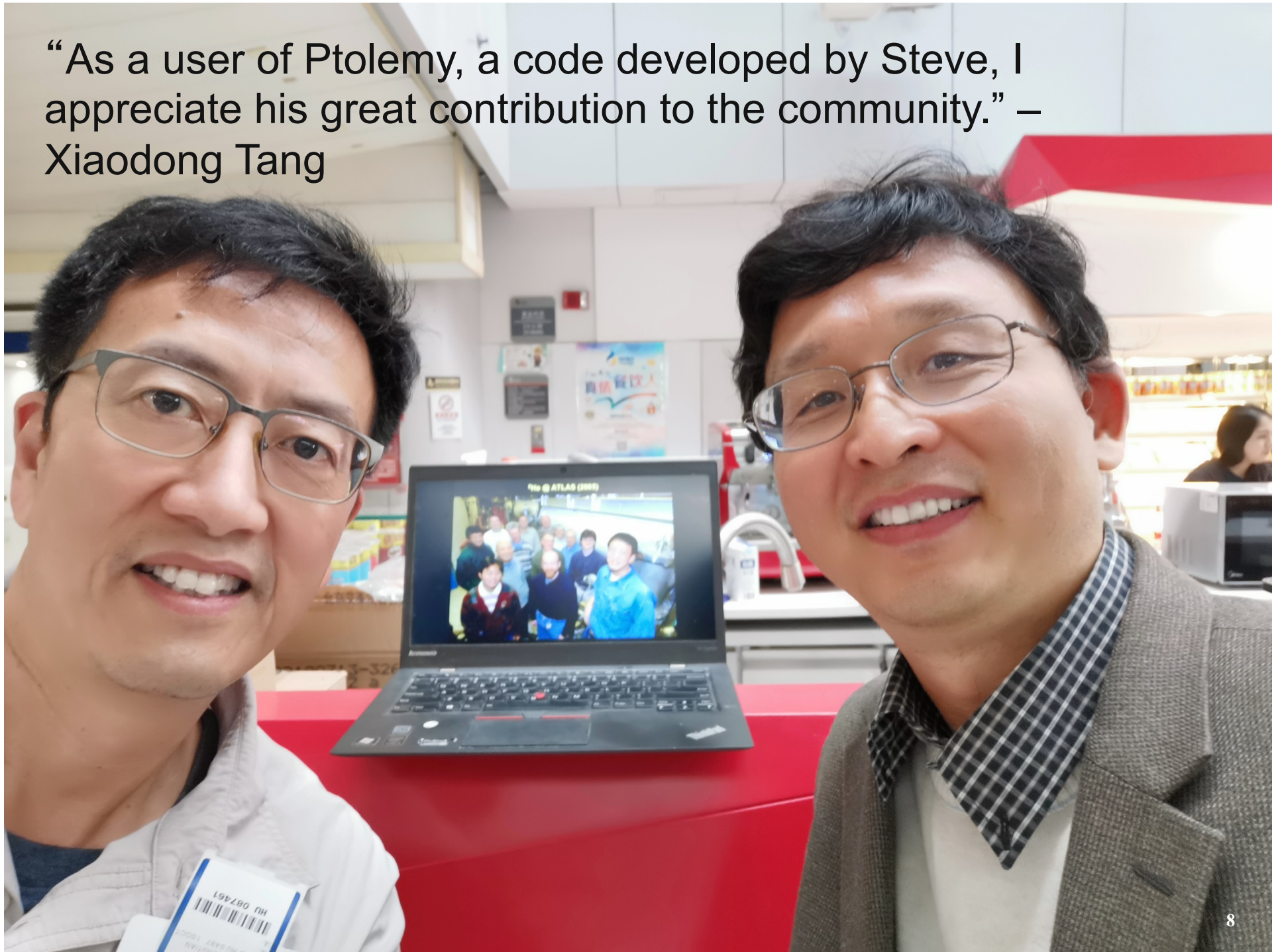
$${}^8\text{He} - {}^4\text{He} : \delta\nu_{8,4} = 64702.509(2) \text{ MHz} - 1.011 (\langle r^2 \rangle_{c,8} - \langle r^2 \rangle_{c,4}) \text{ MHz/fm}^2$$

100 kHz error in IS \rightarrow ~ 1% error in radius



${}^6\text{He}$ @ ATLAS (2005)

“As a user of Ptolemy, a code developed by Steve, I appreciate his great contribution to the community.” – Xiaodong Tang



He-6 Collaboration

P. Mueller, L.-B. Wang, K. Bailey, J.P. Greene, D. Henderson, R.J. Holt, R. Janssens, C.L. Jiang,
Z.-T. Lu, T.P. O'Conner, R.C. Pardo, K.E. Rehm, J.P. Schiffer, X.D. Tang - *Physics, Argonne*
G. W. F. Drake - *Univ of Windsor, Canada*

He-8 Collaboration

P. Mueller, K. Bailey, R. J. Holt, R. V. F. Janssens, Z.-T. Lu, T. P. O'Connor, I. Sulai - *Physics, Argonne*; M.-
G. Saint Laurent, *J.-Ch. Thomas, A.C.C. Villari - GANIL, Caen, France*
G. W. F. Drake - *Univ of Windsor, Canada* L.-B. Wang – *Los Alamos Lab*



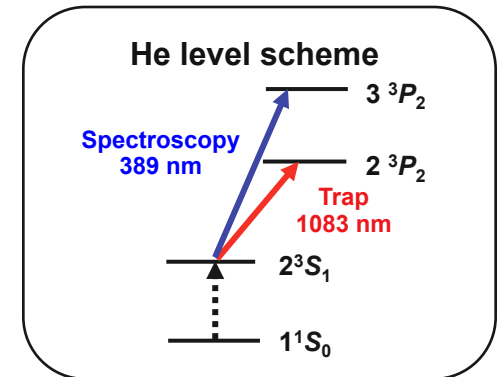
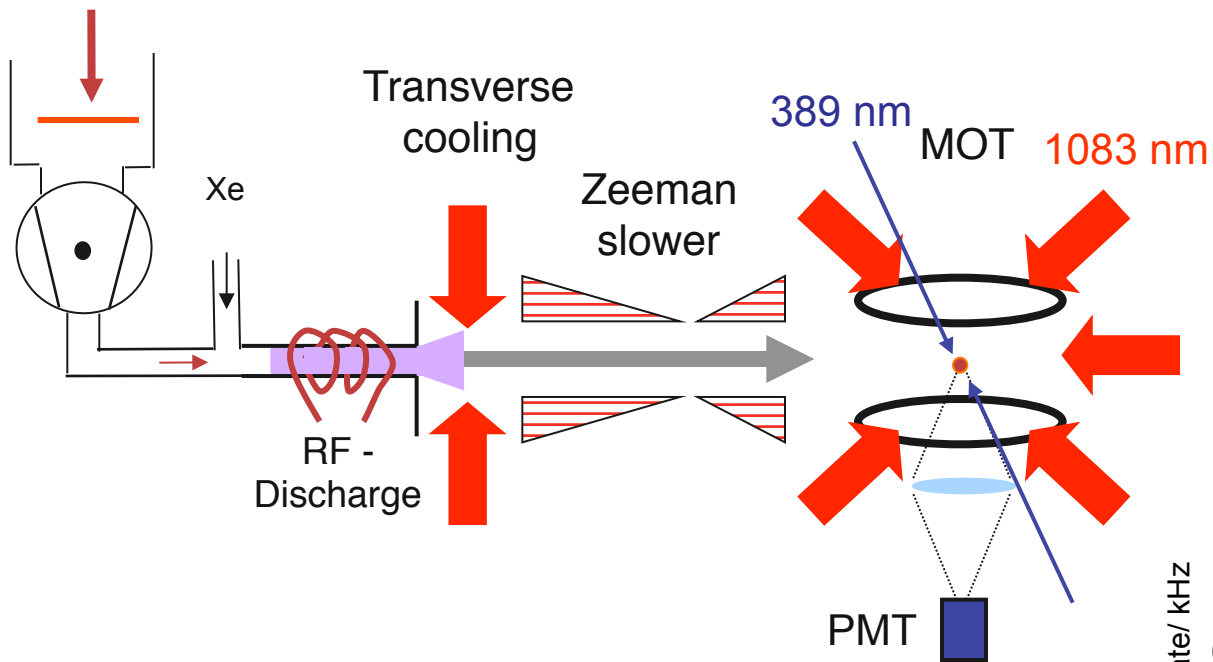
Jan. 26th 2007



Atom Trapping of ${}^6\text{He}$ & ${}^8\text{He}$ at GANIL (2007)

$\sim 1 \times 10^8$ ${}^6\text{He}^+/\text{s}$
 $\sim 5 \times 10^5$ ${}^8\text{He}^+/\text{s}$

Atom Trap Setup

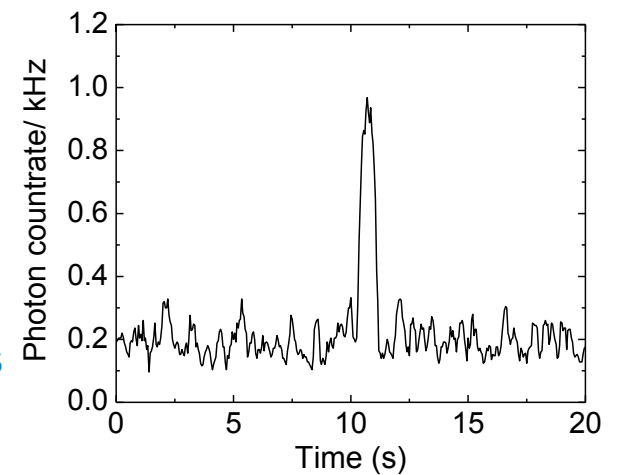


Source
 ${}^6\text{He} \sim 5 \times 10^7/\text{s}$
 ${}^8\text{He} \sim 1 \times 10^5/\text{s}$

Capture efficiency
 1×10^{-7}

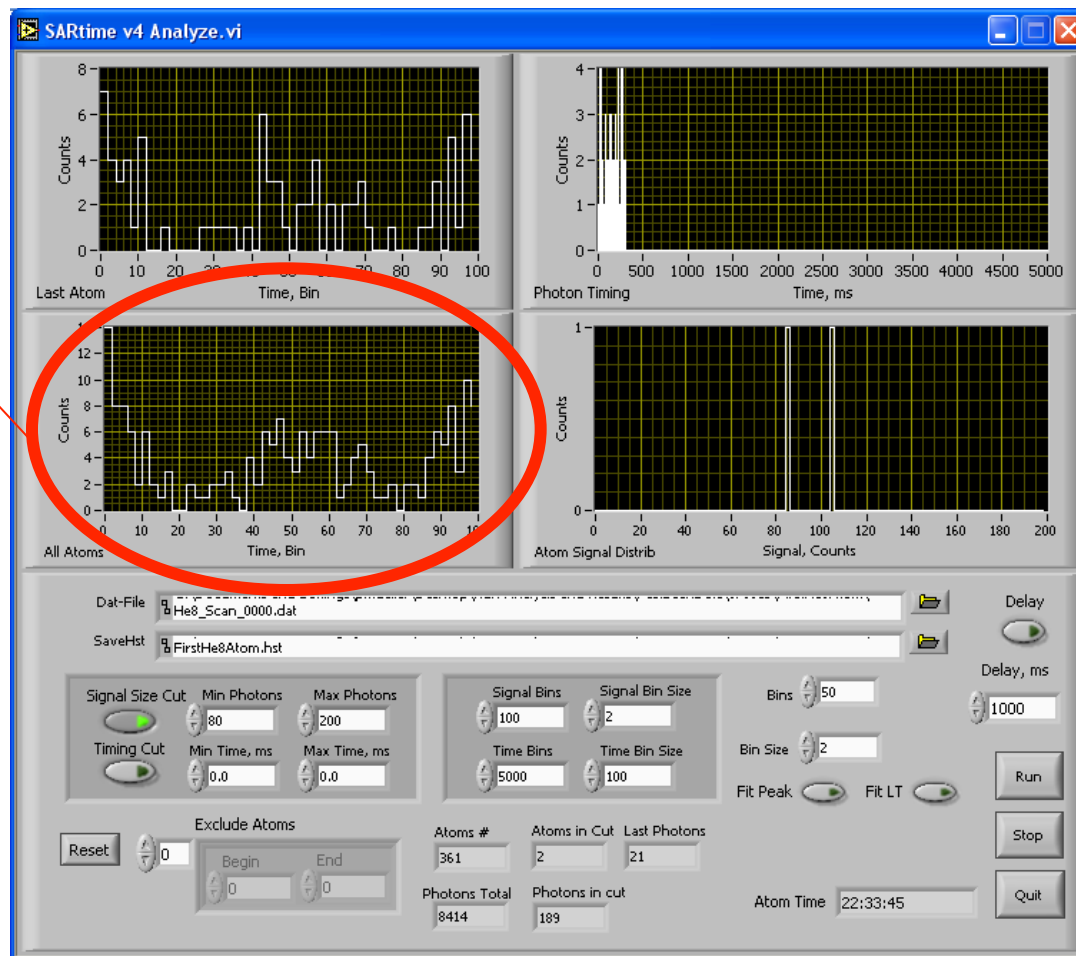
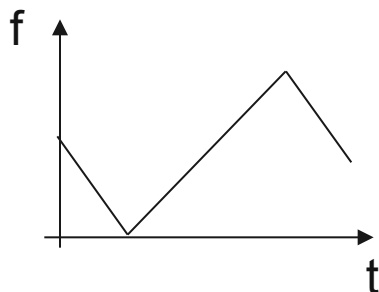
Trap
 ${}^6\text{He} \sim 5/\text{s}$
 ${}^8\text{He} \sim 1 \times 10^{-2}/\text{s}$

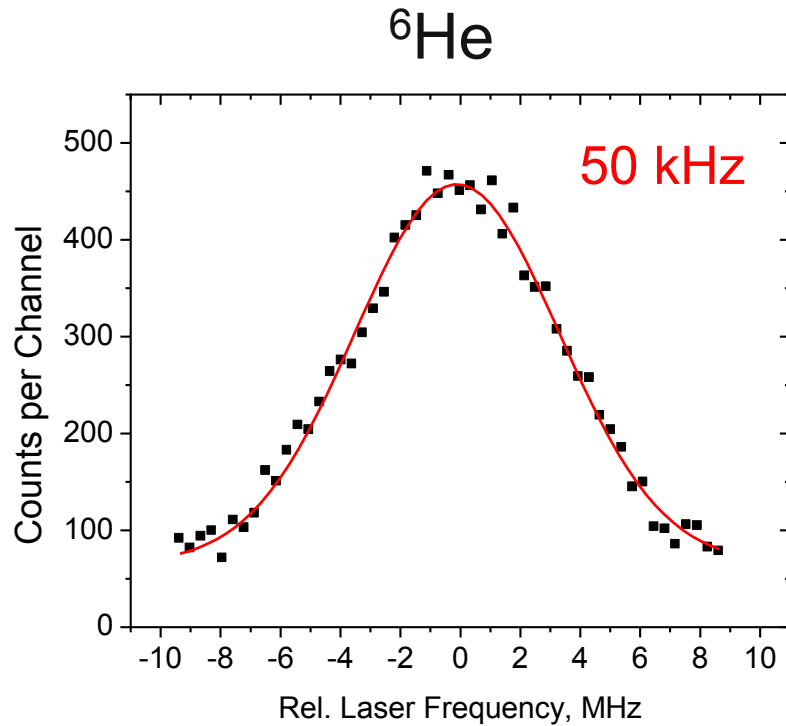
One trapped ${}^6\text{He}$ atom



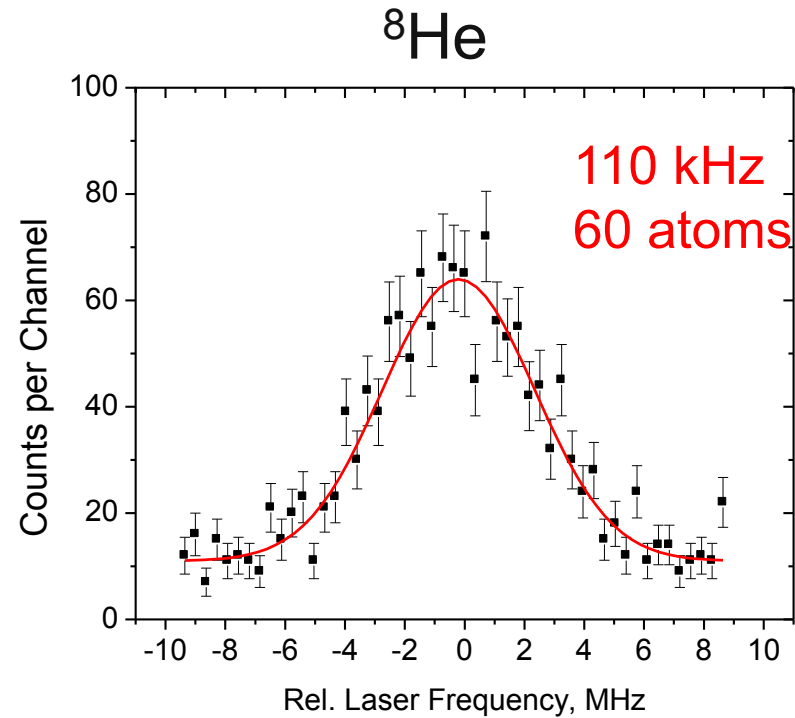
He-8 Trapped!

First He-8 Atom
June 15th 2007





~30 ${}^6\text{He}$ atoms/s



~30 ${}^8\text{He}$ atoms/hr

For $2^3\text{S}_1 - 3^3\text{P}_2$ transition @ 389 nm:

$${}^6\text{He} - {}^4\text{He} : \delta\nu_{6,4} = 43196.171(2) \text{ MHz} - 1.010 (\langle r^2 \rangle_{c,6} - \langle r^2 \rangle_{c,4}) \text{ MHz/fm}^2$$

$${}^8\text{He} - {}^4\text{He} : \delta\nu_{8,4} = 64702.509(2) \text{ MHz} - 1.011 (\langle r^2 \rangle_{c,8} - \langle r^2 \rangle_{c,4}) \text{ MHz/fm}^2$$

Field Shifts:

$${}^6\text{He} - {}^4\text{He} : \delta\nu_{6,4} = 1.430 (7) [30] \text{ MHz}$$

$${}^8\text{He} - {}^4\text{He} : \delta\nu_{8,4} = 1.020 (42) [45] \text{ MHz}$$

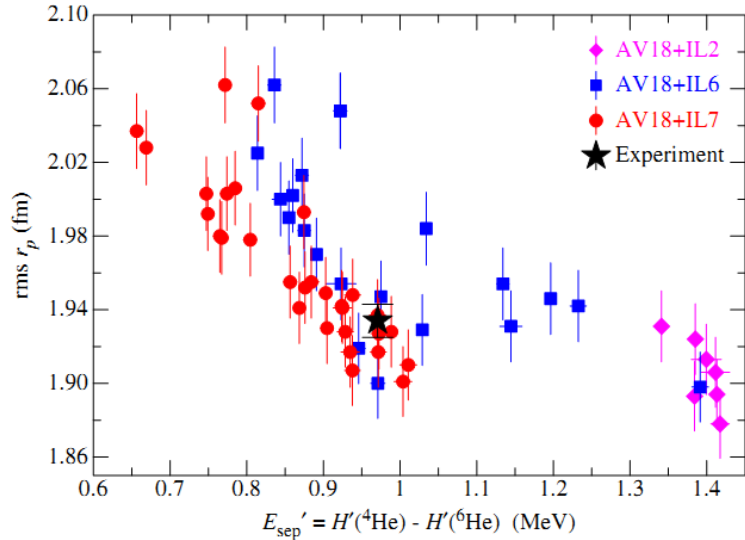
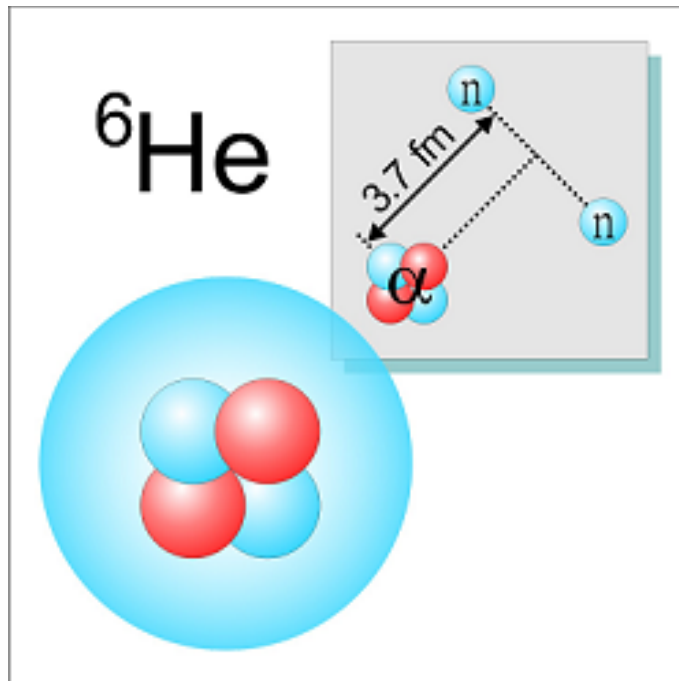


FIG. 9 (color online). ${}^6\text{He}$ point-proton radius vs two-neutron separation energy obtained from a number of GFMC calculations with different initial conditions and three-body potentials.

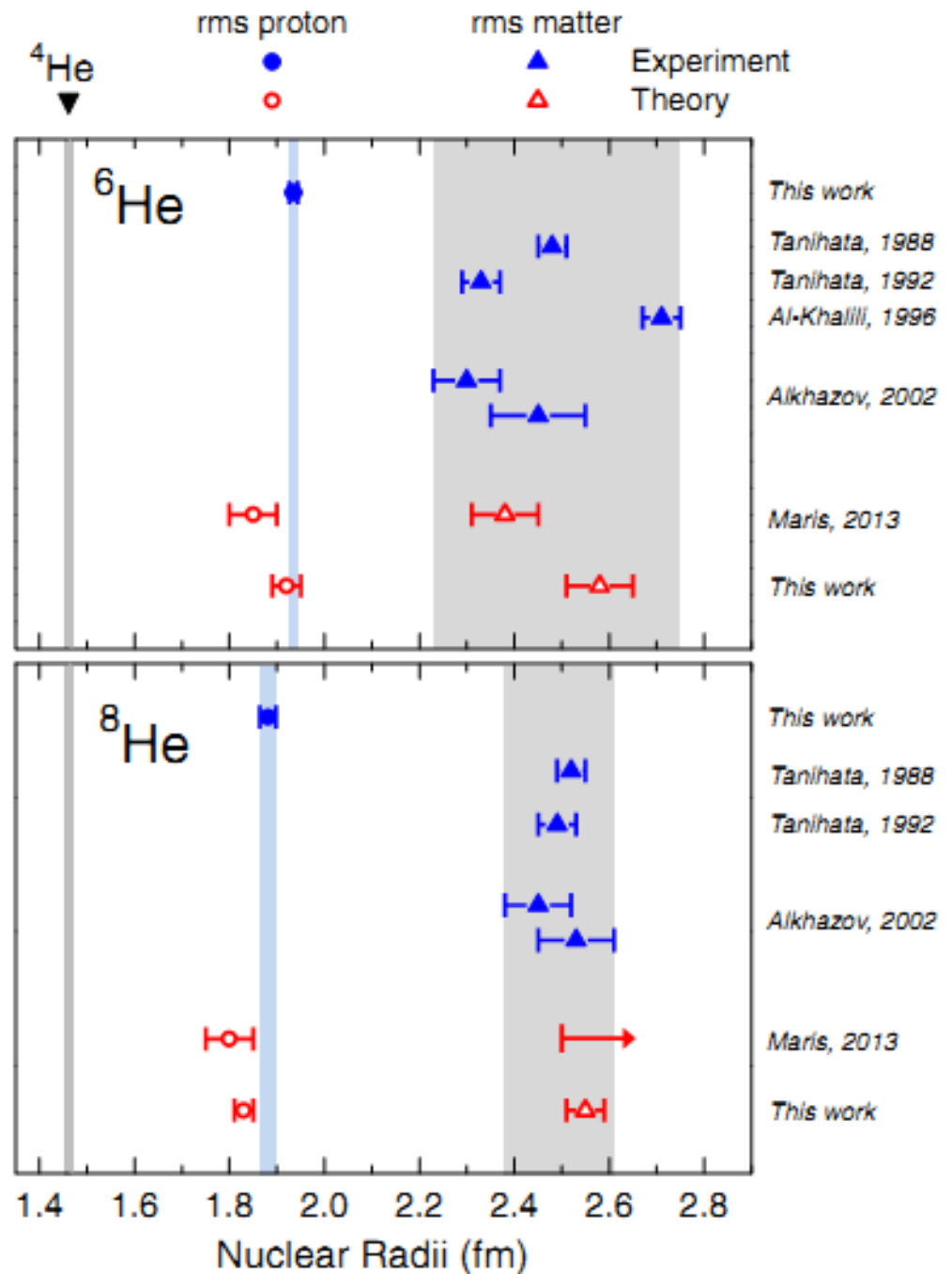
Steve Pieper wrote in RMP (2013)

GFMC calculations using AV18 with the Illinois V_{ijk} are successful in reproducing the energies of nuclear states for $A \leq 12$ (Pieper, Wiringa, and Carlson, 2004; Pieper, 2005, 2008b). The He isotope energies and corresponding two-neutron separation energies E_{2n} obtained for AV18 + IL7 are given in Table V. Quantities other than energy may converge at a much slower pace. This is particularly true for the radii of weakly bound states. For ${}^8\text{He}$, and especially for ${}^6\text{He}$, there are long-term fluctuations in the radii as GFMC calculations propagate in imaginary time. These fluctuations are associated with the small two-neutron separation energies; according to the calculations, $E_{\text{sep}} = 0.97$ MeV for ${}^6\text{He}$ and 1.86 MeV for ${}^8\text{He}$. Figure 9 displays the results of multiple GFMC calculations with different initial conditions. Even though GFMC can compute the binding energies precisely, its relative errors in E_{sep} are significant since, for ${}^6\text{He}$, E_{sep} is only 3% of the binding energy. For example, changes in the starting wave function Ψ_T and other aspects of the GFMC calculations can result in changes of 0.2 MeV in E_{sep} , or a few percent change in the radius. For these weakly bound nuclei, more precise values of radii can be obtained by selecting those calculations that simultaneously yield the experimentally known E_{sep} value, marked with a star in Fig. 9, with its associated range interpreted as an uncertainty of the computed radii. The same procedure is used to get the matter radii r_m (Table VI). The computed point-proton radii

Proton- & Matter-Radii of ${}^6\text{He}$ & ${}^8\text{He}$



Lu *et al.*, RMP (2013)





Atom Trap Trace Analysis – Method and Applications in the Earth Sciences

Zheng-Tian Lu

Department of Modern Physics
University of Science and Technology of China (USTC)

Hefei National Laboratory for Physical Sciences at the Microscale

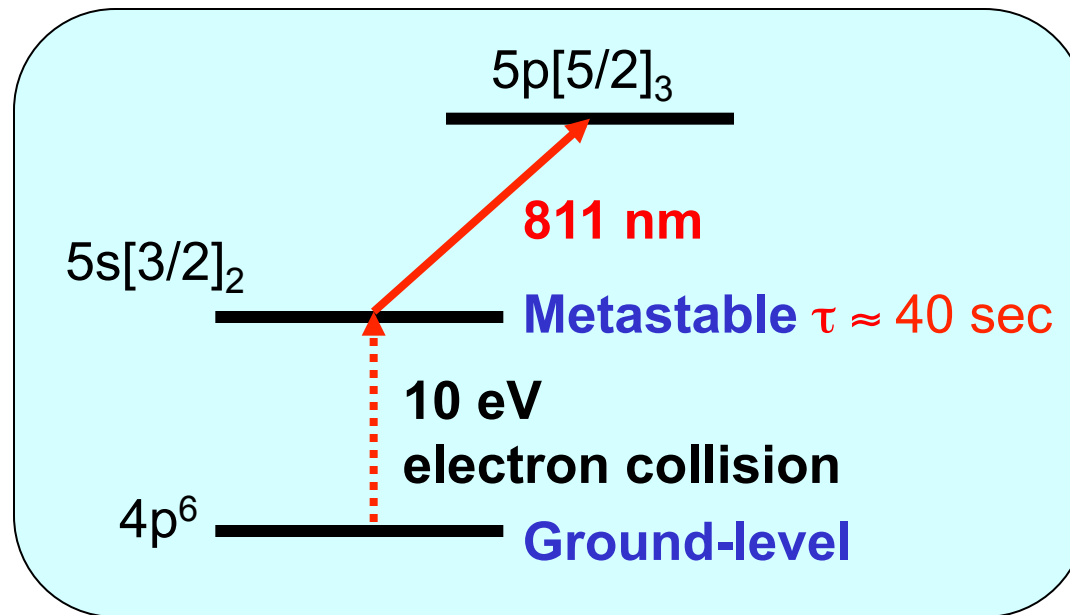
CAS Center for Excellence in Quantum Information
and Quantum Physics

ICEQT September 2019

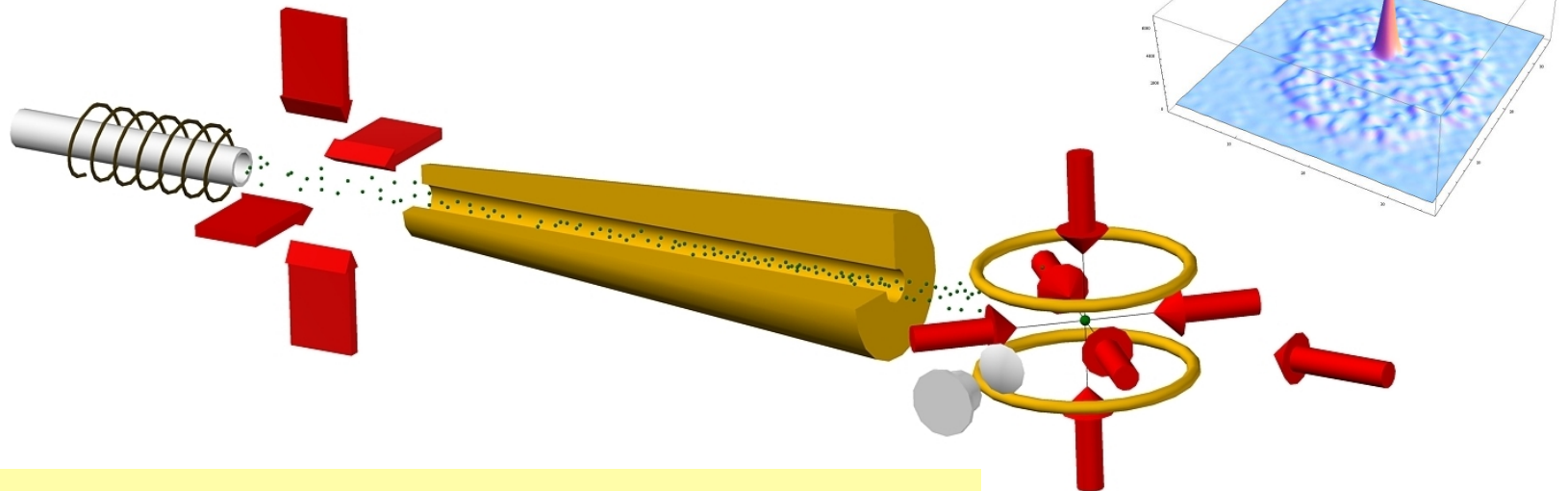
創寰宇學府
育天下英才

嚴濟慈題
一九八八年五月

Atom Trap Trace Analysis (ATTA) - Krypton



CCD image profile of a single ^{81}Kr atom

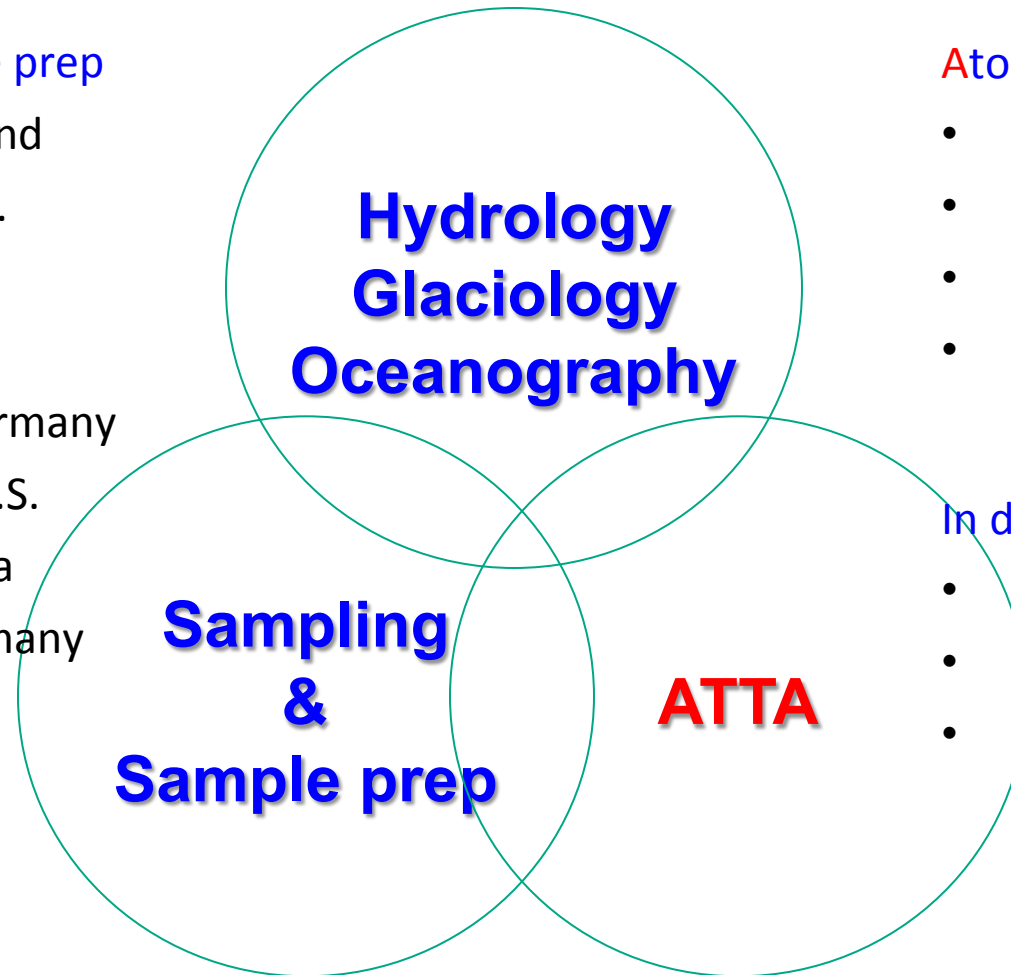


Ultrasensitive isotope trace analyses with a magneto-optical trap
Chen *et al.*, Science 286, 1139 (1999)

Collaboration of Earth Science and Physics

Sampling & sample prep

- Bern, Switzerland
- U. Chicago, U.S.
- IAEA, Vienna
- USTC, China
- Heidelberg, Germany
- U. Delaware, U.S.
- CSIRO, Australia
- Hamburg, Germany
-



Atom Trap Trace Analysis

- Argonne, U.S.
- Heidelberg, Germany
- USTC, China
- CSIRO, Australia

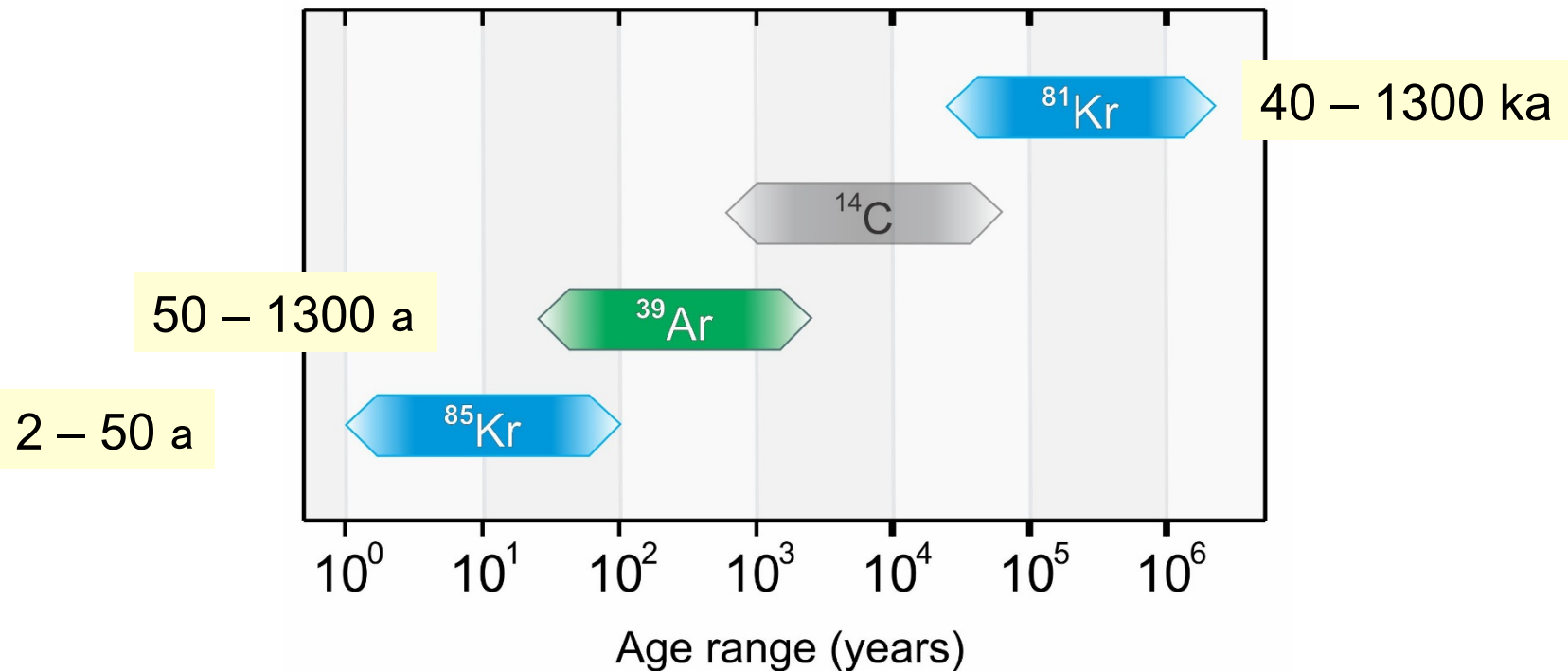
In development

- IAEA, UN
- Hamburg, Germany
- Bern, Switzerland

Google "ATTA Primer"

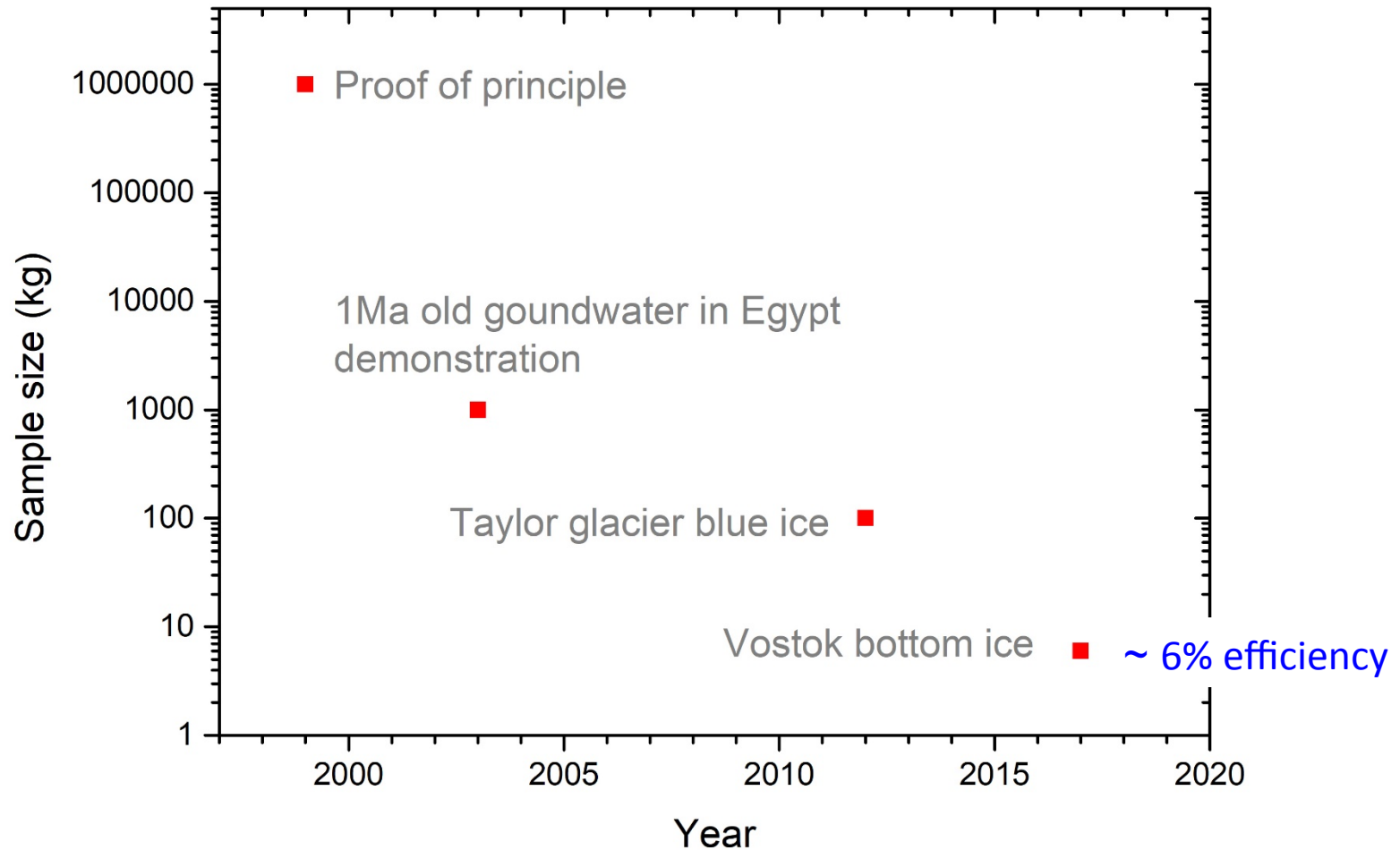
Radioactive gas isotopes for tracing and dating

- **Gas:** Even distribution around the world
- **Inert:** Simple production, transport and deposition process

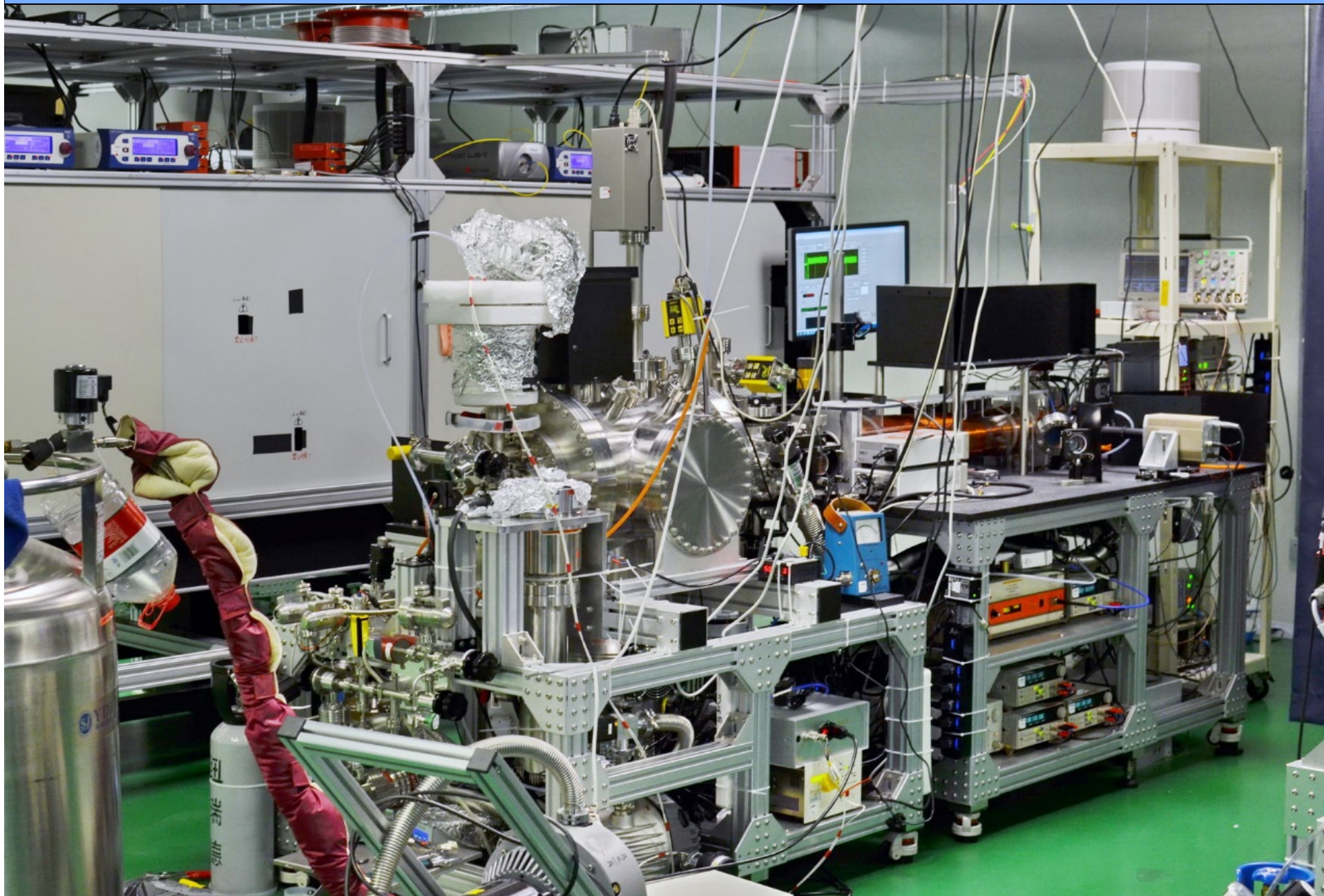


^{85}Kr , ^{39}Ar , ^{81}Kr : Ideal isotopes for dating / tracing

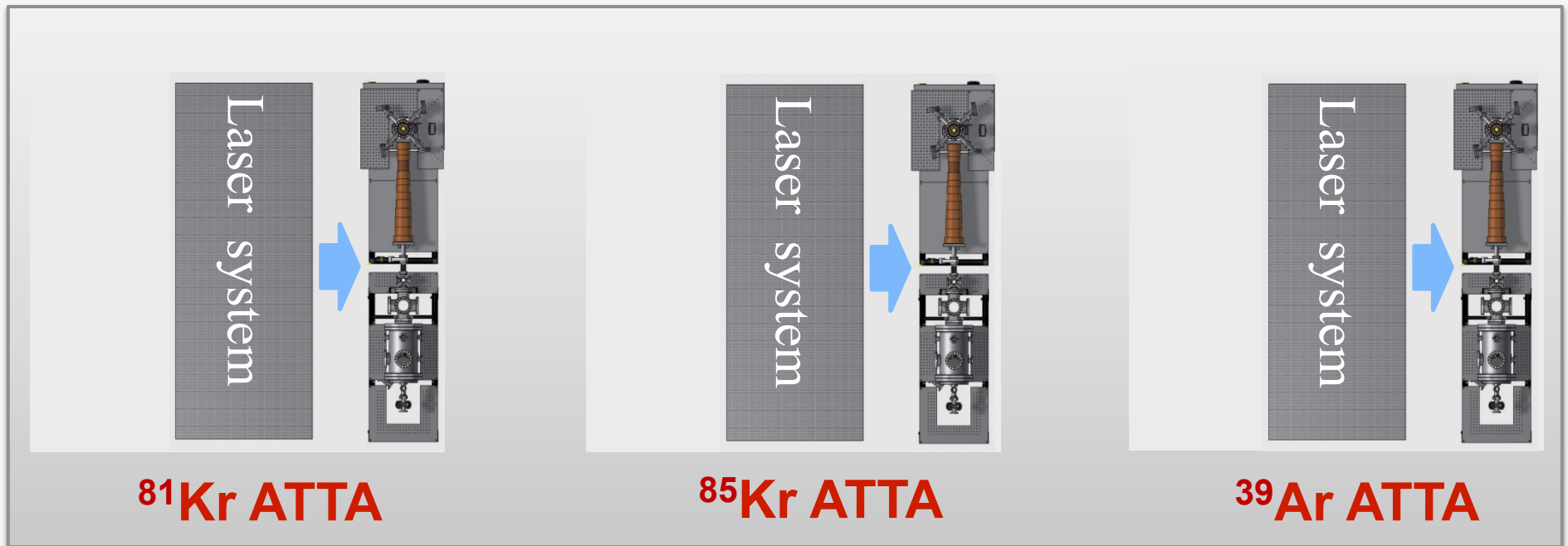
Water/Ice sample size needed over the years



ATTA Instrument for Kr-81 at USTC

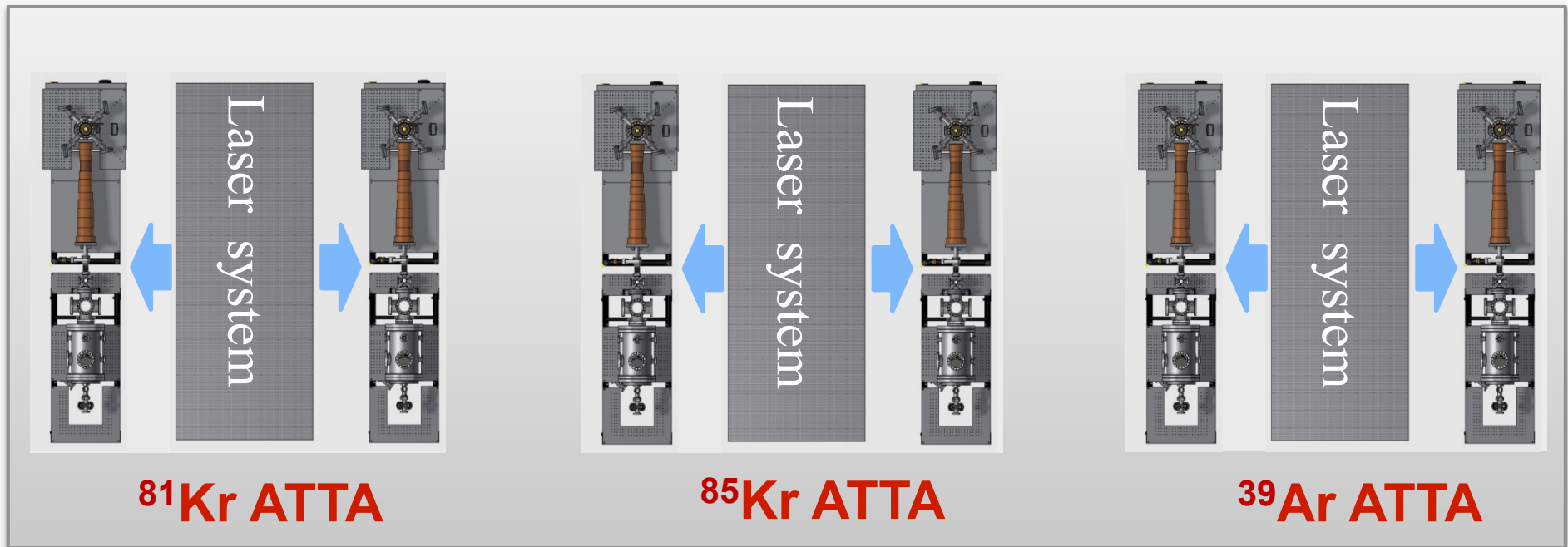


USTC ATTA Lab (2019)



- 1x Kr-81 beamlines, 200 samples/year
- 1x Kr-85 beamlines, 500 samples/year
- 1x Ar-39 beamlines, 200 samples/year

USTC ATTA Lab (2021)

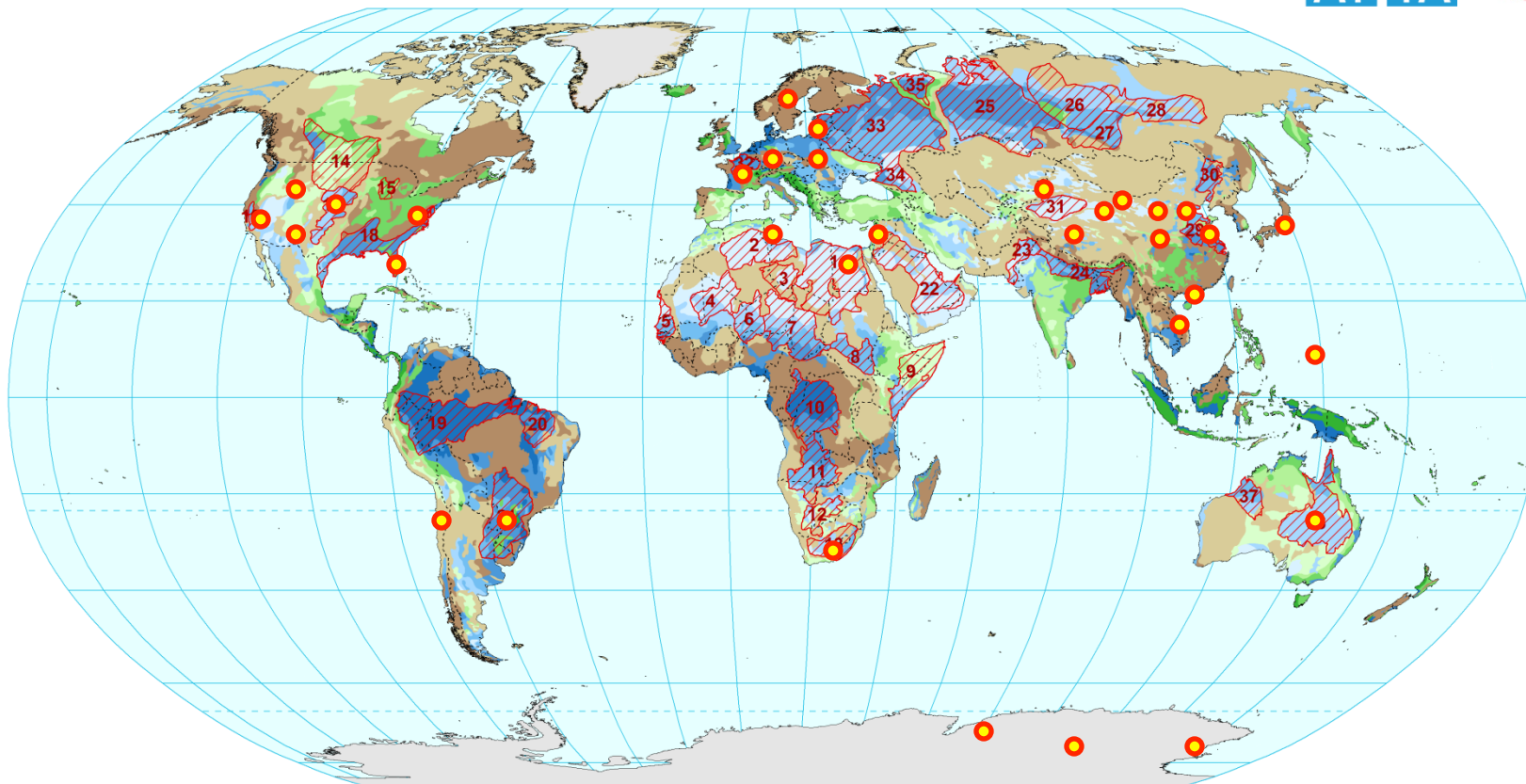


- 2x Kr-81 beamlines, 400 samples/year
- 2x Kr-85 beamlines, 1000 samples/year
- 2x Ar-39 beamlines, 400 samples/year



Groundwater Resources of the World

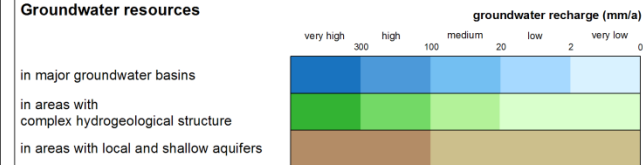
- Large Aquifer Systems -



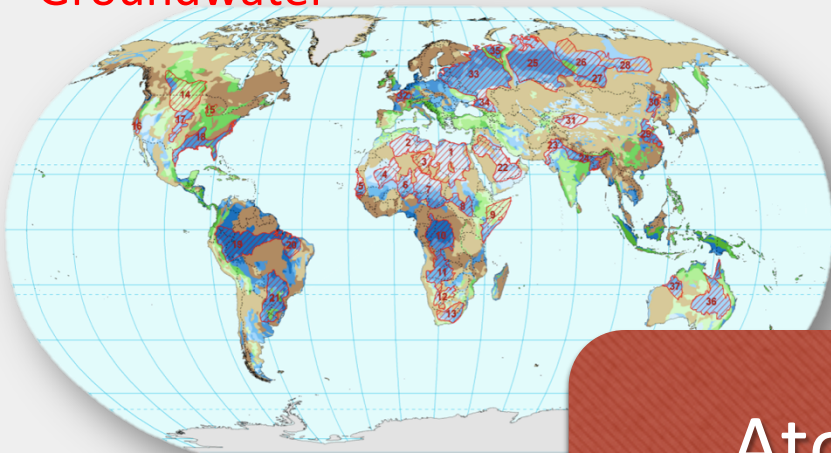
Large Aquifer Systems

- | | | | |
|--------------------------------------------|-----------------------------------------------------|--------------------------------------|----------------------------------|
| 1. Nubian Aquifer System (NAS) | 11. Northern Kalahari Basin | 21. Guarani Aquifer System | 31. Tarim Basin |
| 2. Northwest Sahara Aquifer System (NWSAS) | 12. Southeast Kalahari Basin | 22. Arabian Aquifer System | 32. Parisian Basin |
| 3. Murzuk-Djado Basin | 13. Karoo Basin | 23. Indus Basin | 33. East European Aquifer System |
| 4. Taoudeni-Tanezrouft Basin | 14. Northern Great Plains / Interior Plains Aquifer | 24. Ganges-Brahmaputra Basin | 34. North Caucasus Basin |
| 5. Senegalo-Mauritanian Basin | 15. Cambro-Ordovician Aquifer System | 25. West Siberian Artesian Basin | 35. Pechora Basin |
| 6. Iullemeden-Irhazer Aquifer System | 16. California Central Valley Aquifer System | 26. Tunguss Basin | 36. Great Artesian Basin |
| 7. Chad Basin | 17. High Plains-Ogallala Aquifer | 27. Angara-Lena Artesian Basin | 37. Canning Basin |
| 8. Sudd Basin (Umm Ruwaba Aquifer) | 18. Gulf Coastal Plains Aquifer System | 28. Yakut Basin | |
| 9. Ogaden-Juba Basin | 19. Amazonas Basin | 29. North China Plain Aquifer System | |
| 10. Congo Intracratonic Basin | 20. Maranhao Basin | 30. Songliao Basin | |

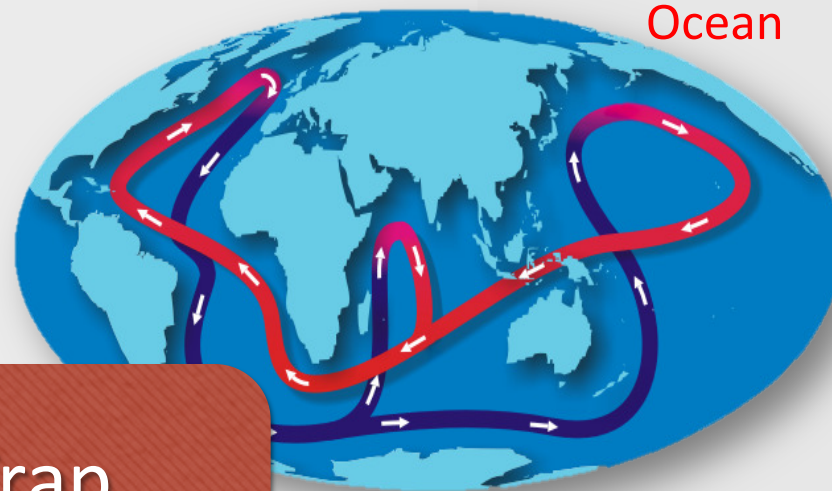
Groundwater resources



Groundwater



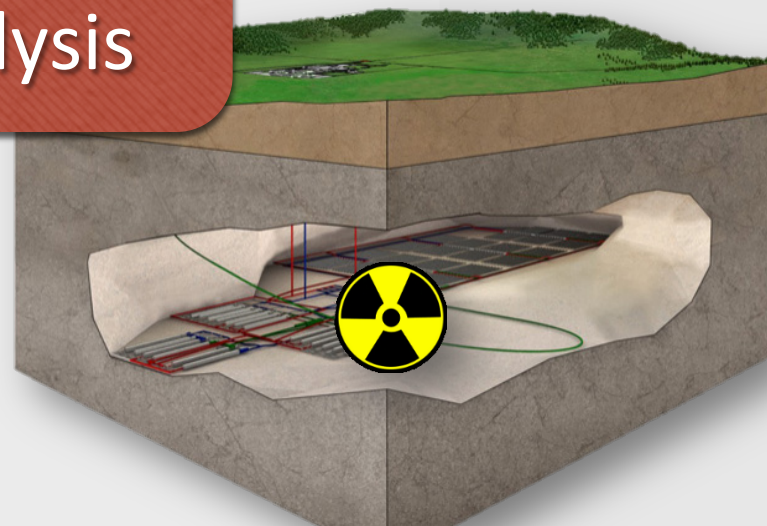
Ocean



Atom Trap Trace Analysis



Glacier



Nuclear safety



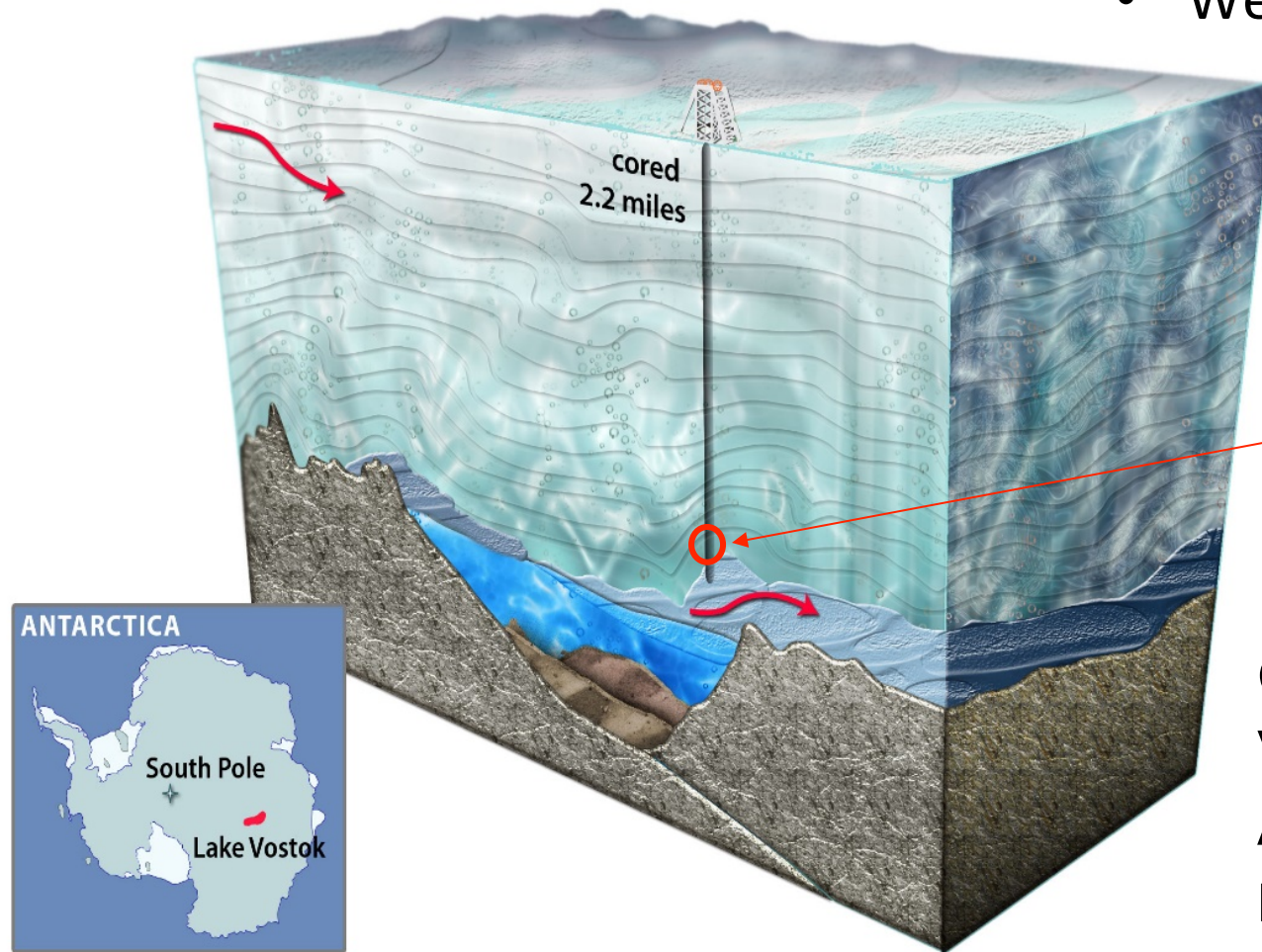


- ^{81}Kr characterize old (isolated) groundwater environment for sites of nuclear waste repository
- Waste Isolation Pilot Plant (WIPP), New Mexico, USA
 - China National Nuclear Corp., Beishan, China
 - Central Res. Inst. of Electric Power Industry (CRIEPI), Japan

Beishan, China

Vostok Ice Core

- Depth : Below 3500m
- 3 Consecutive samples
- Weight : ~ **6 kg!**

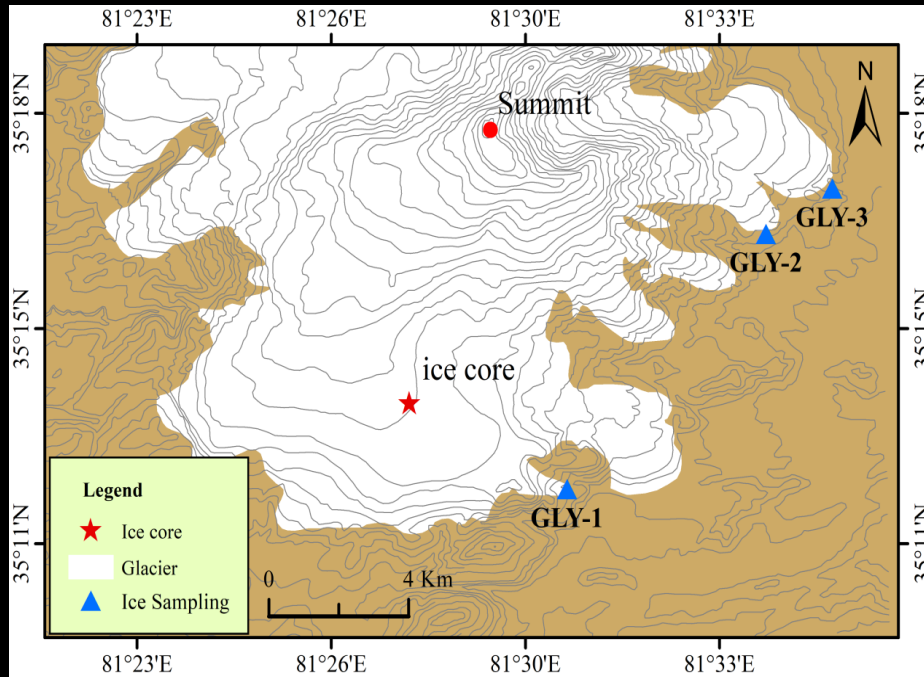


1 million years old

Collaborators:
Vladimir Lipenkov, AARI
Amaelle Landais, LSCE
Barbara Stenni, Venice

With Laboratory for Sciences of Climate and Environment (LSCE)
Arctic and Antarctic Research Institute (AARI)

Guliya Ice Cap



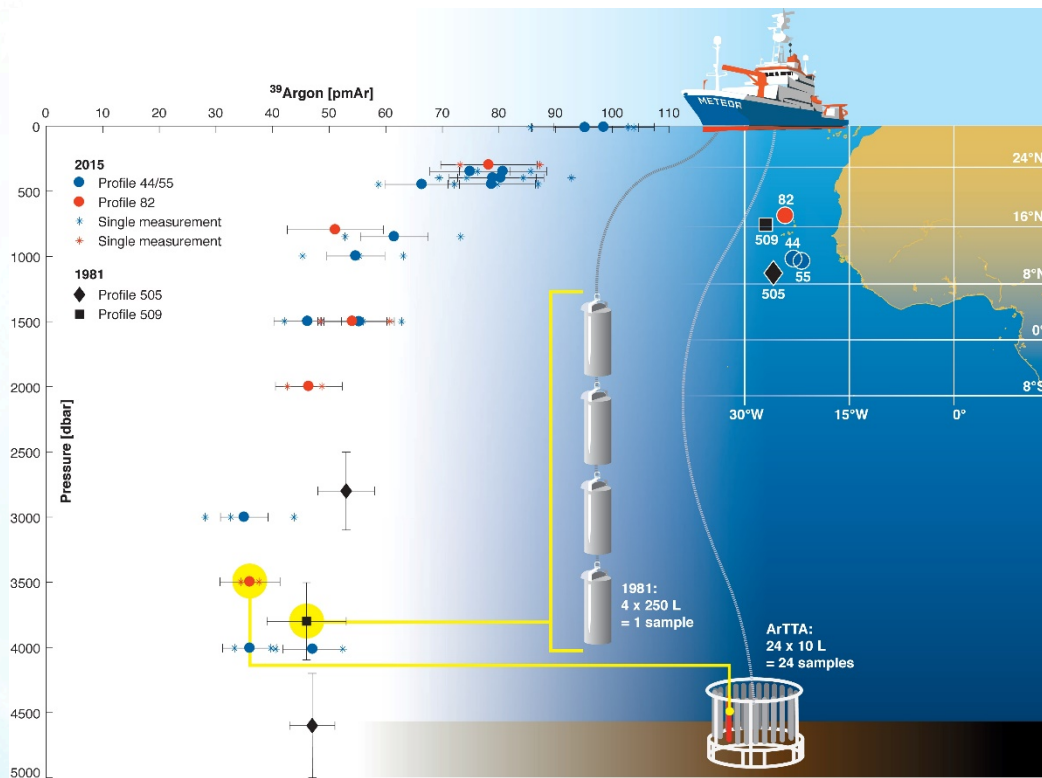
^{81}Kr ages of 8 samples
Ages < 15 - 74 ka

Li-De Tian *et al.*,
Geophys. Res. Lett. (2019)



First ^{39}Ar -dating of small ocean samples

Heidelberg University (Markus Oberthaler and Werner Aeschbach)
GEOMAR Kiel (Toste Tanhua)



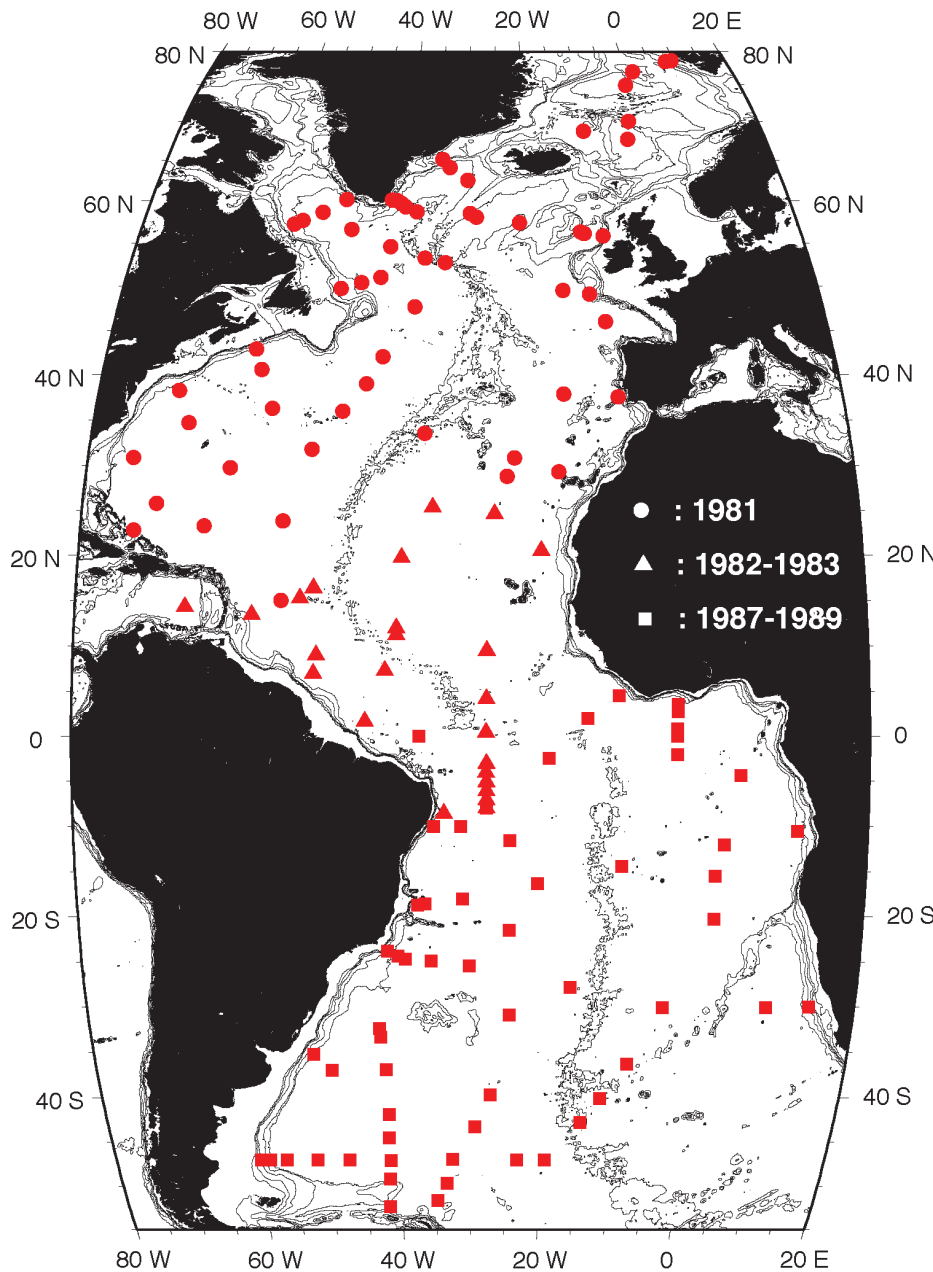
Ebser, S. et al., *Nat. Commun.* **9**, 5046 (2018)

Sampling:

- Three depth profiles from the North Atlantic (Eastern tropical North Atlantic Oxygen Minimum Zone)
- **5 liters of water** per data point
- Sampling with **standard Niskin bottles**

Main results:

- ^{39}Ar & CFC constrain transit time distributions
- Mean ages of up to 800 years
- Reveal ocean ventilation patterns
- Advection in intermediate depths much more important than previously assumed

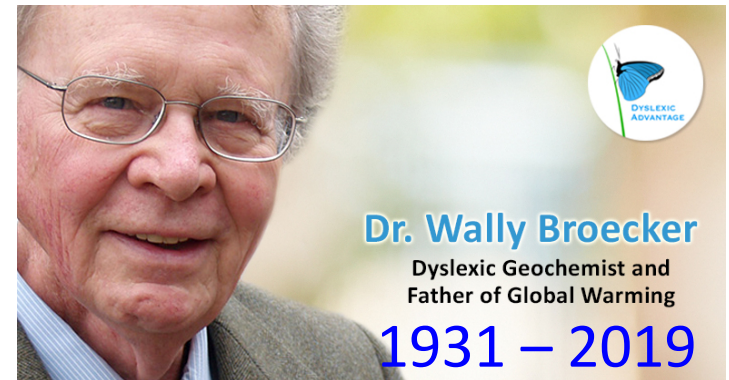


Lamont-Doherty Earth Observatory
 Columbia University
 William Smethie Jr., Martin Stute *et al.*

- 900 Atlantic samples collected in the 80's
- ^{39}Ar dating by USTC and Heidelberg

“a more dense survey of ^{39}Ar with higher accuracy measurements would prove of great value in constraining ocean general circulation models.”

--- Broecker and Peng (2000)



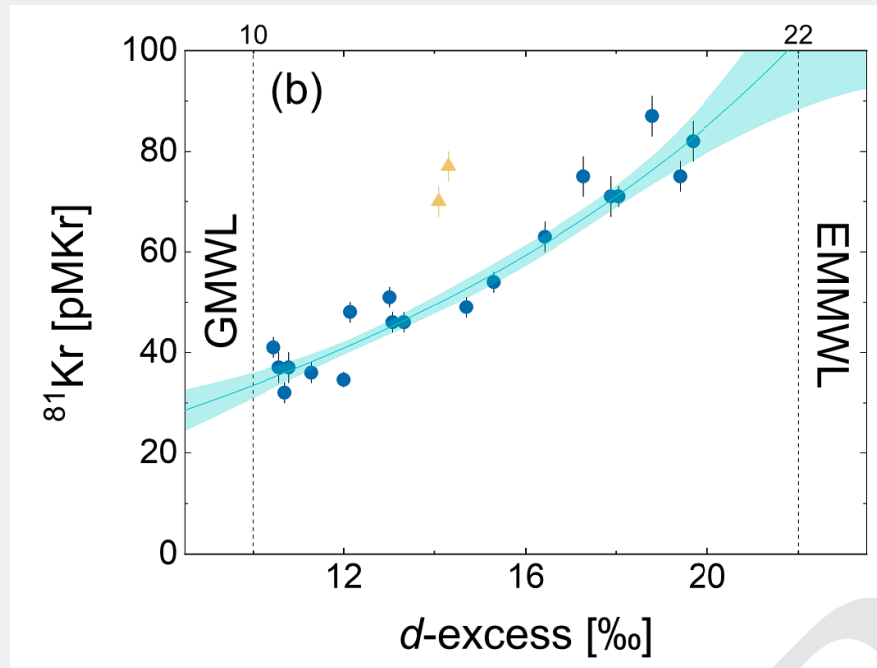
Z.-T. Lu *et al.*, *Earth-Sci. Rev.*
 (2014)



Reika Yokochi
U Chicago

Radiokrypton unveils dual moisture sources of a deep desert aquifer

Yokochi *et al.*, PNAS 116, 16222 (2019)



Atlantic ocean
361±30 ka

Mediterranean
< 38 ka

- Investigated the paleo-hydroclimate properties of the **Nubian Sandstone Aquifer** in the **Negev Desert, Israel**.
- Resolved subsurface mixing and identified two distinct moisture sources of recharge.
- Reveals that **tectonically active terrain** can store groundwater.

August 13, 2019 | vol. 116 | no. 33 | pp. 16153–16656

PNAS

Proceedings of the National Academy of Sciences of the United States of America

www.pnas.org

Groundwater source tracing using radiokrypton



High-pressure methane hydrate phase
Injected fluid diffusion and induced seismicity
Social genetic effects on adolescent smoking
Yeast antiviral pathway and apoptosis

^{81}Kr in the Atmosphere

- production rate and abundance

(A) Geomagnetic field

(B) Production rate

(C) ^{81}Kr age correction

< 4%

-- C. Buizert *et al.*, Proc. Natl. Acad.

Sci. **111**, 6876 (2014)

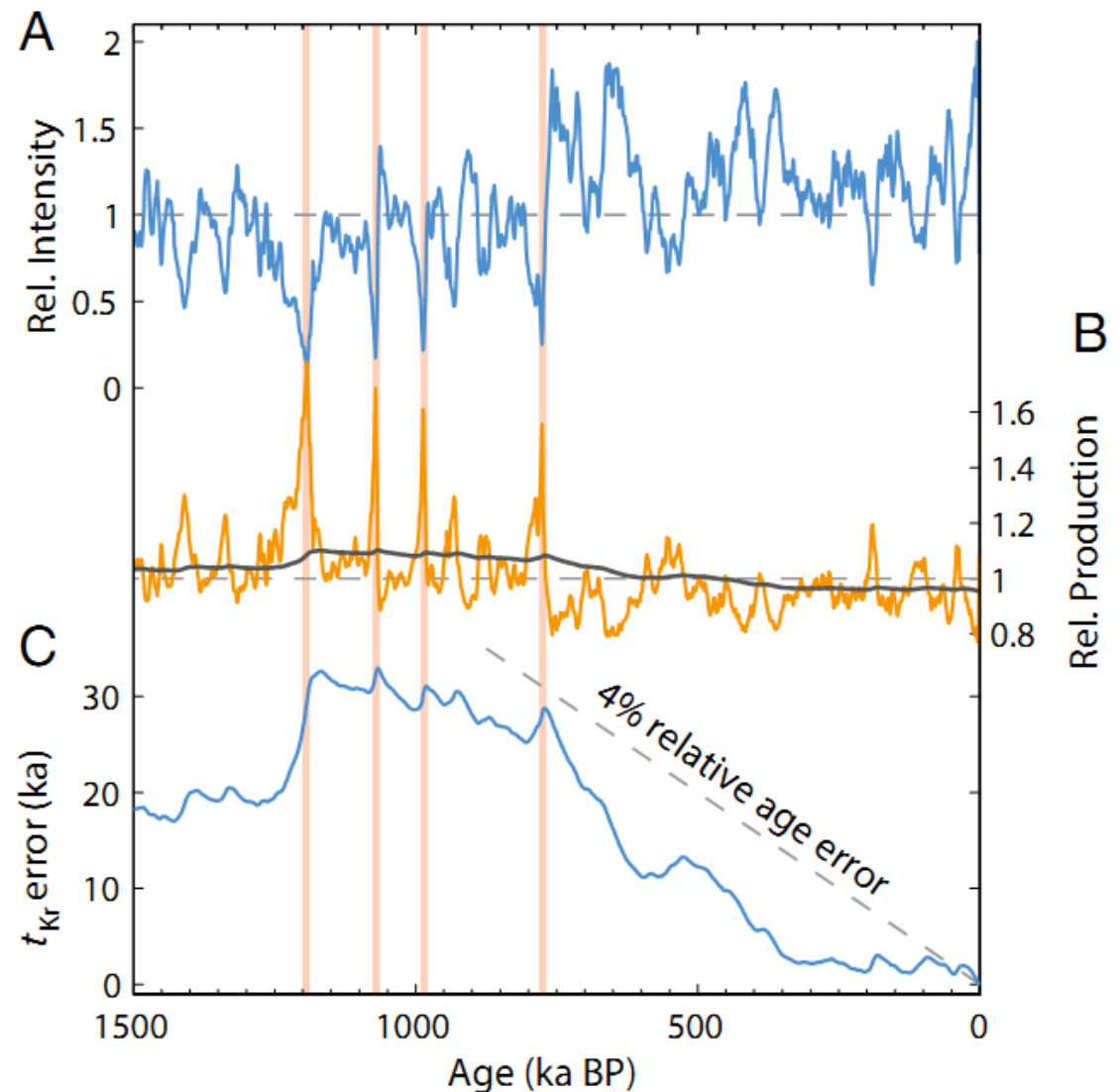


Fig. 3. Stability of atmospheric ^{81}Kr . (A) Relative paleointensity of the geomagnetic field (56). Magnetic reversals are indicated by vertical lines. (B) Relative spallogenic production rate (orange) with relative ^{81}Kr abundance (black). The ^{81}Kr abundance is calculated through a convolution of the production rate with the atmospheric ^{81}Kr residence [$1/330 \times \exp(-t/330)$, with time t in ka]. (C) Estimated error in ^{81}Kr radiometric age when assuming stable atmospheric $^{81}\text{Kr}/\text{Kr}$; positive values indicate an underestimated t_{Kr} .